

Removing Galactic foregrounds on CMB polarization maps using convolutional neural networks

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Co-advisor: Dra. Cora Dvorkin



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A brief summary about me

- PhD in physics student at UNSAM, advised by Daniel Supanitsky
- Co-advised by Cora Dvorkin (Harvard University Professor)
- Member of QUBIC collaboration
- Member of CMB-S4 collaboration
- I spent six months working with Cora at Harvard University (from January-July)

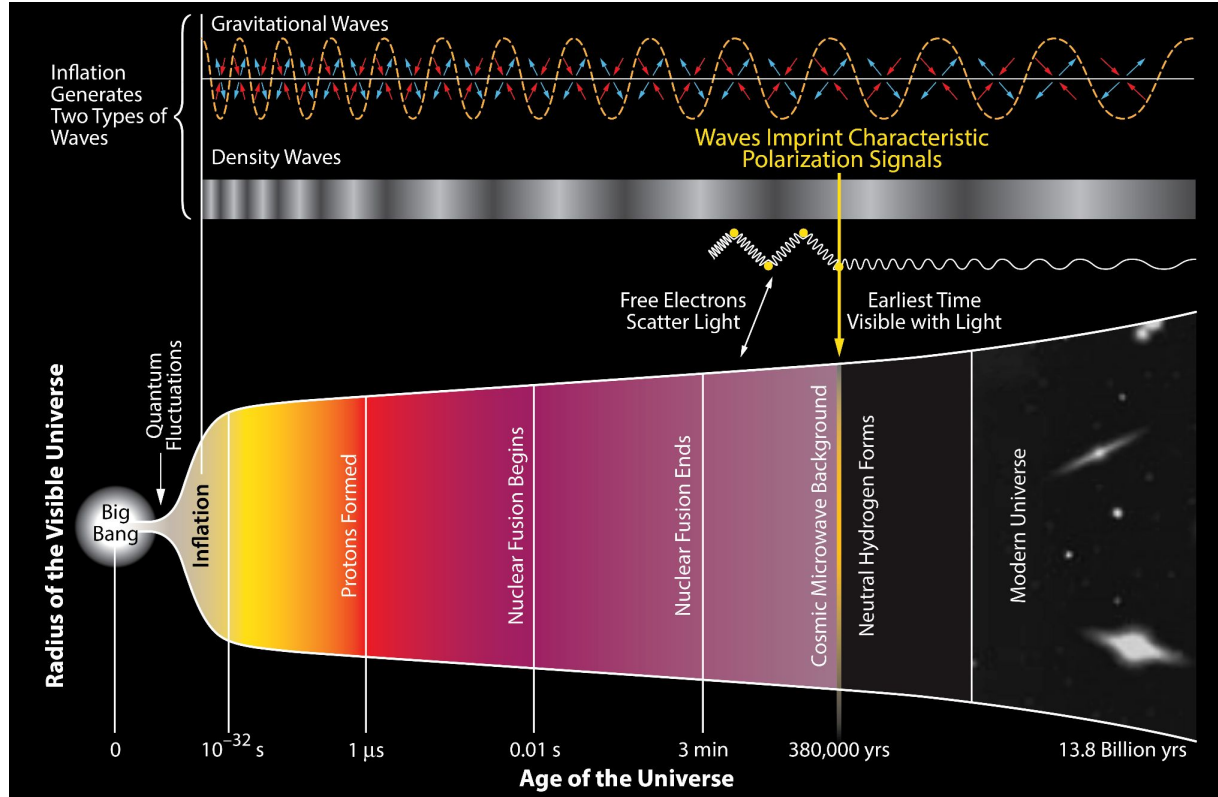


A brief summary about me

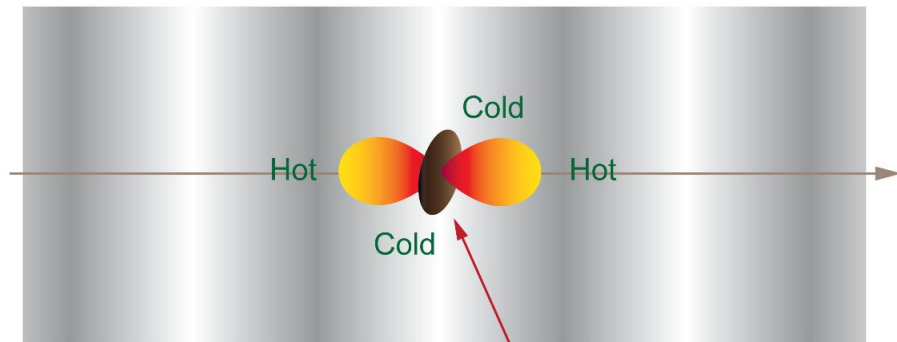
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- Member of QUBIC collaboration
- Member of CMB-S4 collaboration
- Field hockey player & coach
- Huge fan of River Plate (futbol)



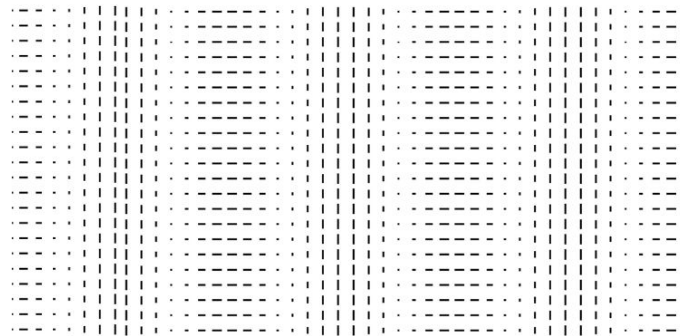
Primordial waves in the Universe



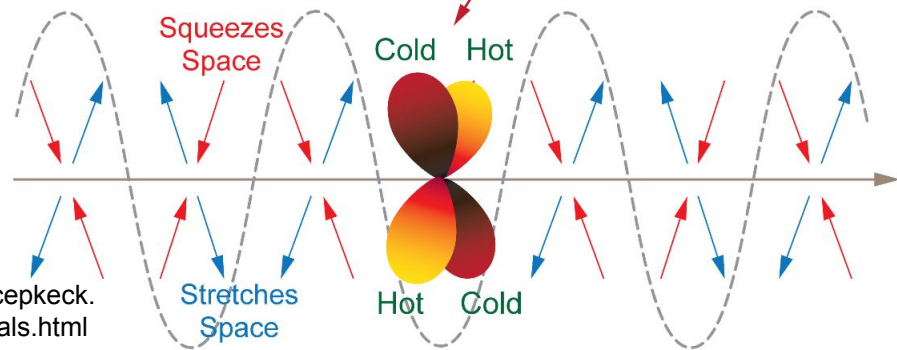
Density Wave



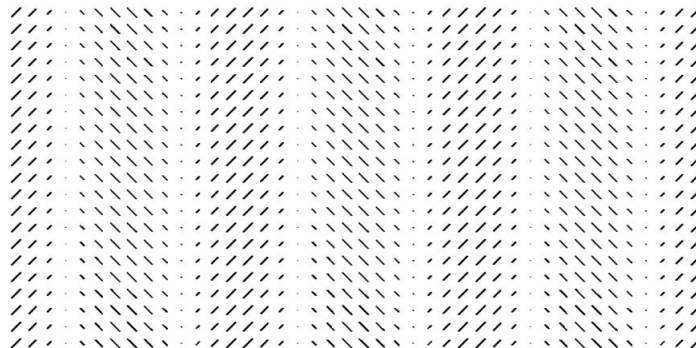
E-Mode Polarization Pattern



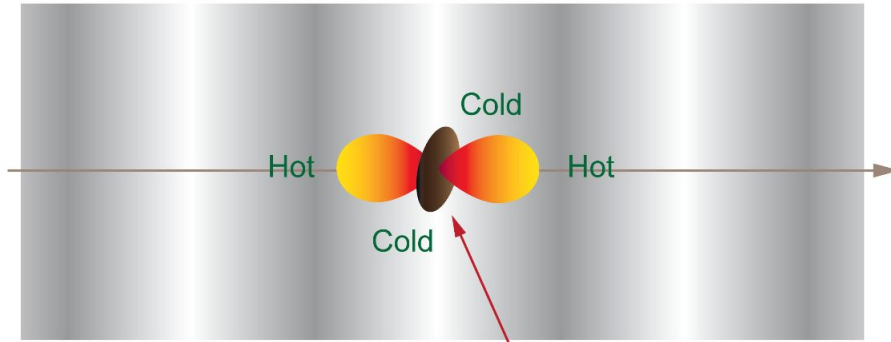
Gravitational Wave



B-Mode Polarization Pattern

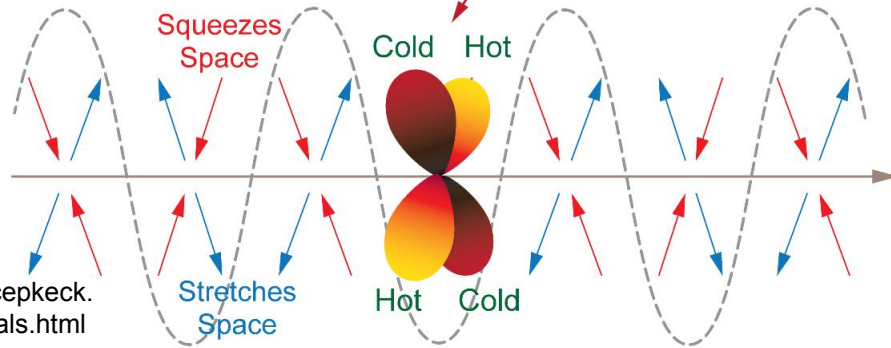


Density Wave

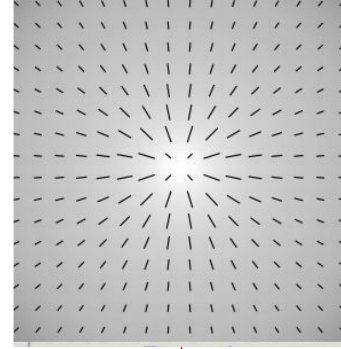


Temperature
Pattern Seen
by Electrons

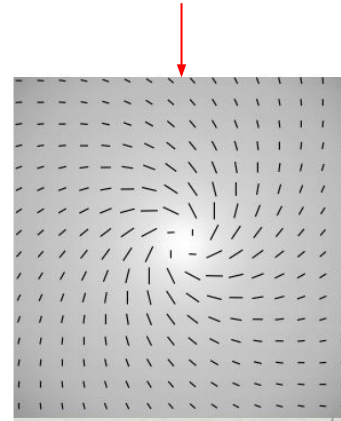
Gravitational Wave



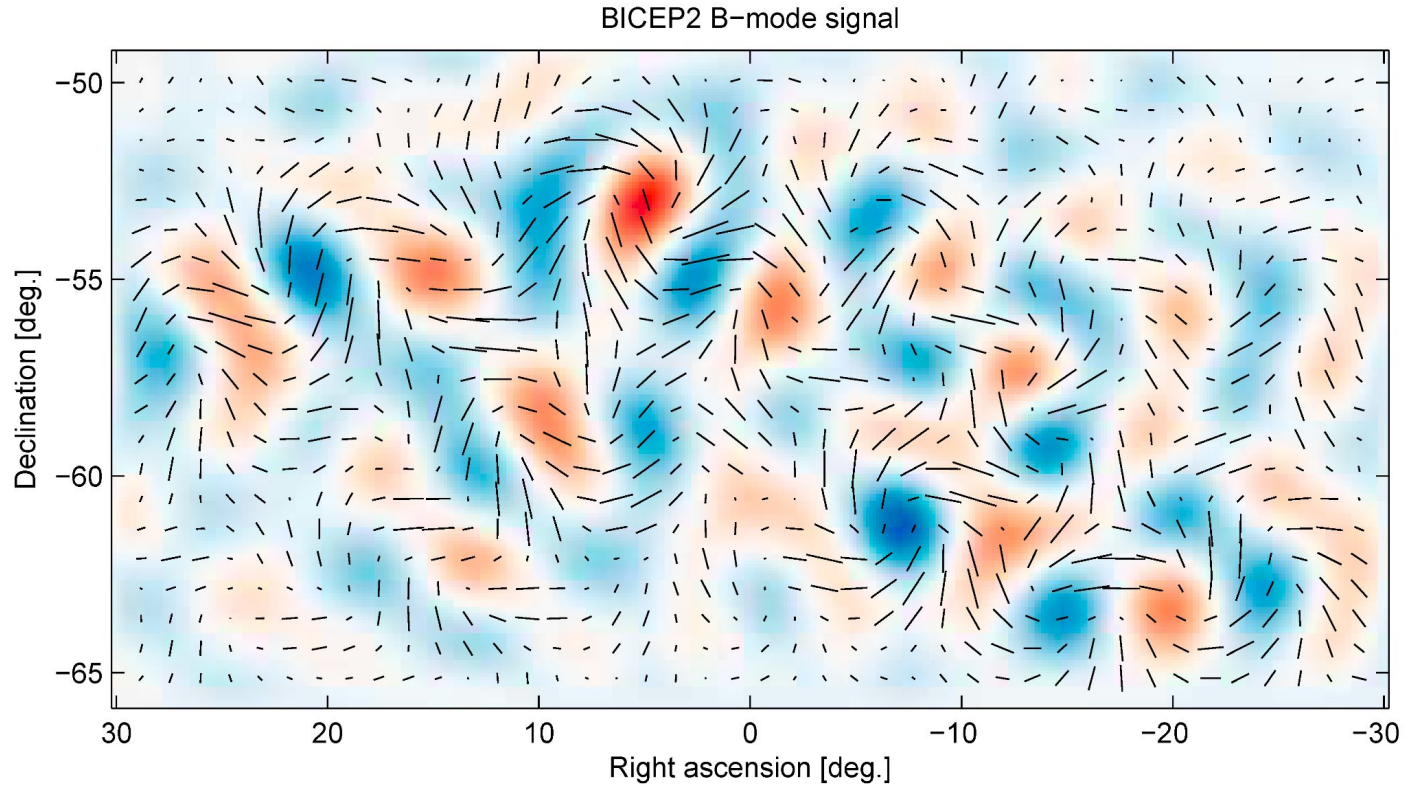
<http://bicepkeck.org/visuals.html>



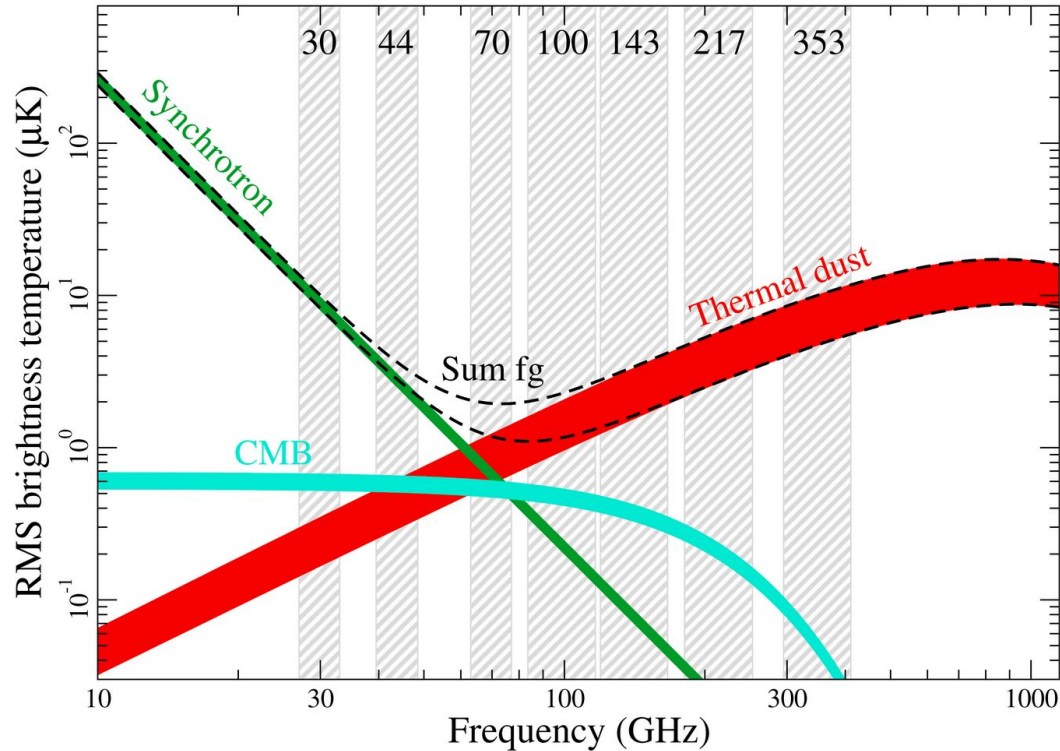
Adding several Fourier modes we obtain sky patterns like these



The best measurement so far

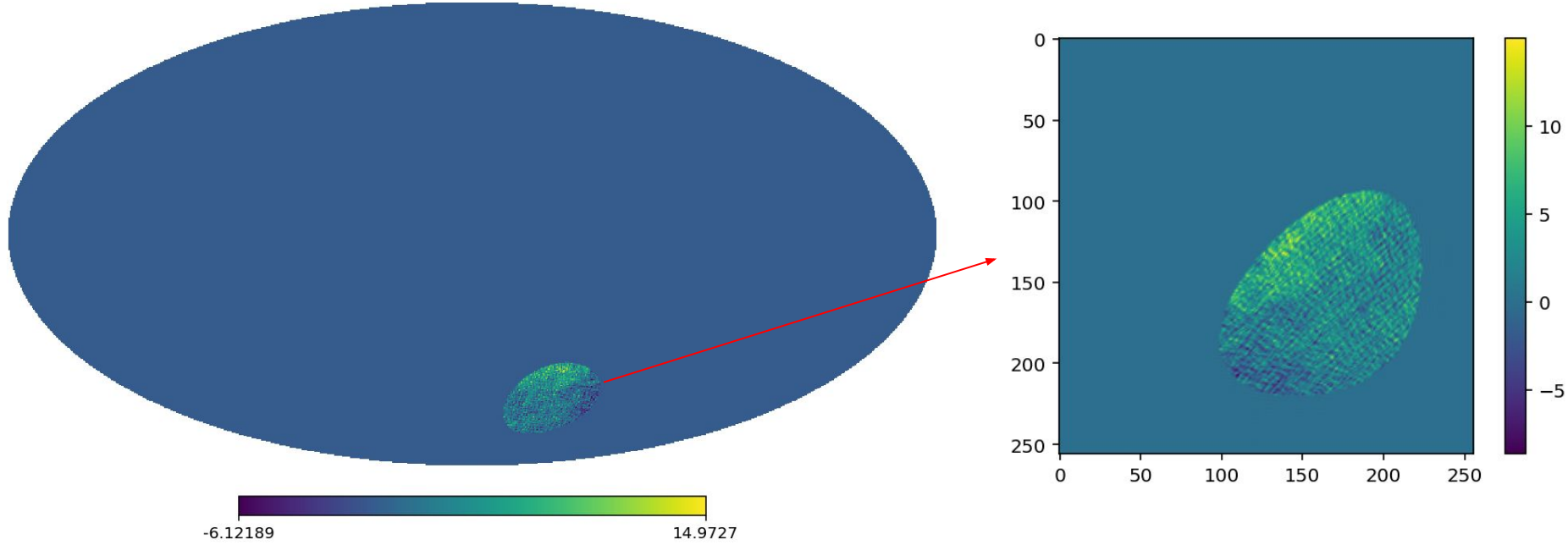


The Galactic foregrounds contamination

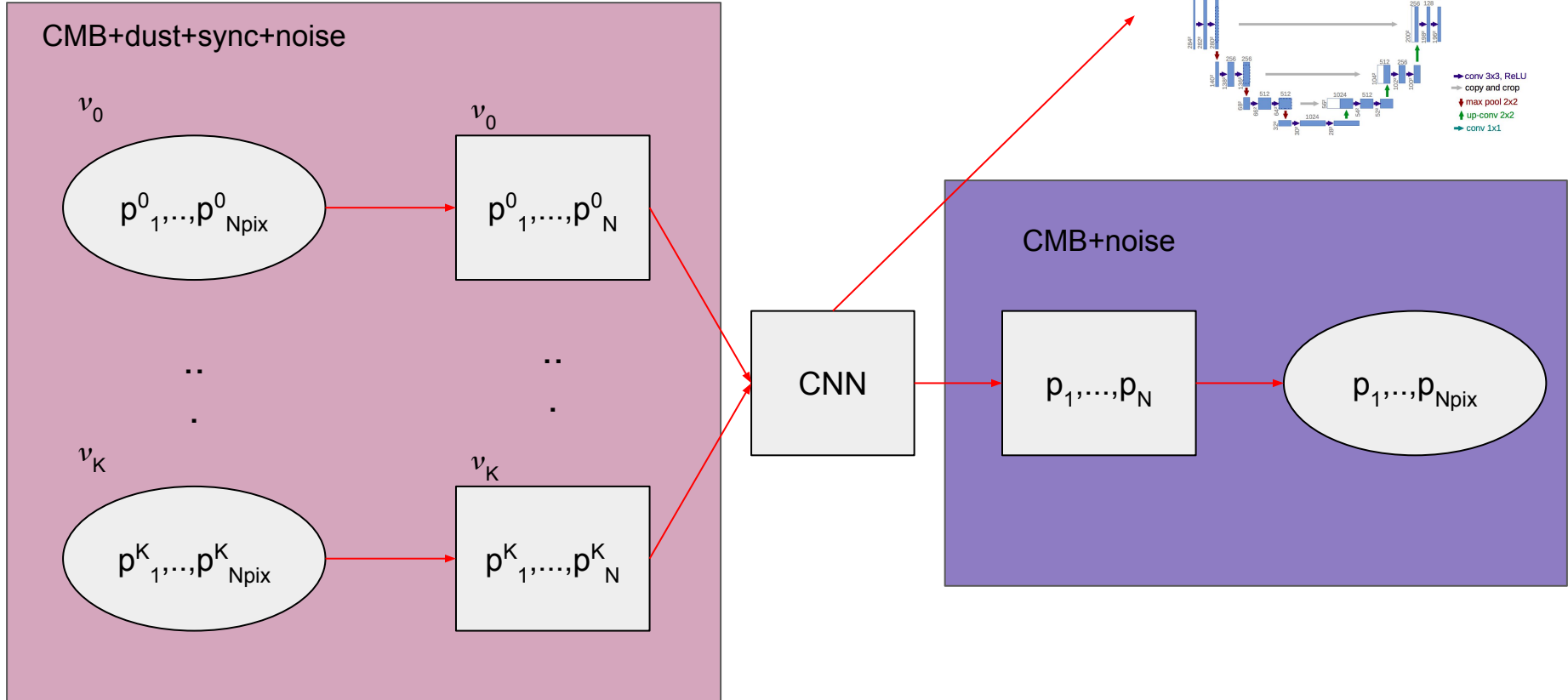


Our current research line

Mollweide view



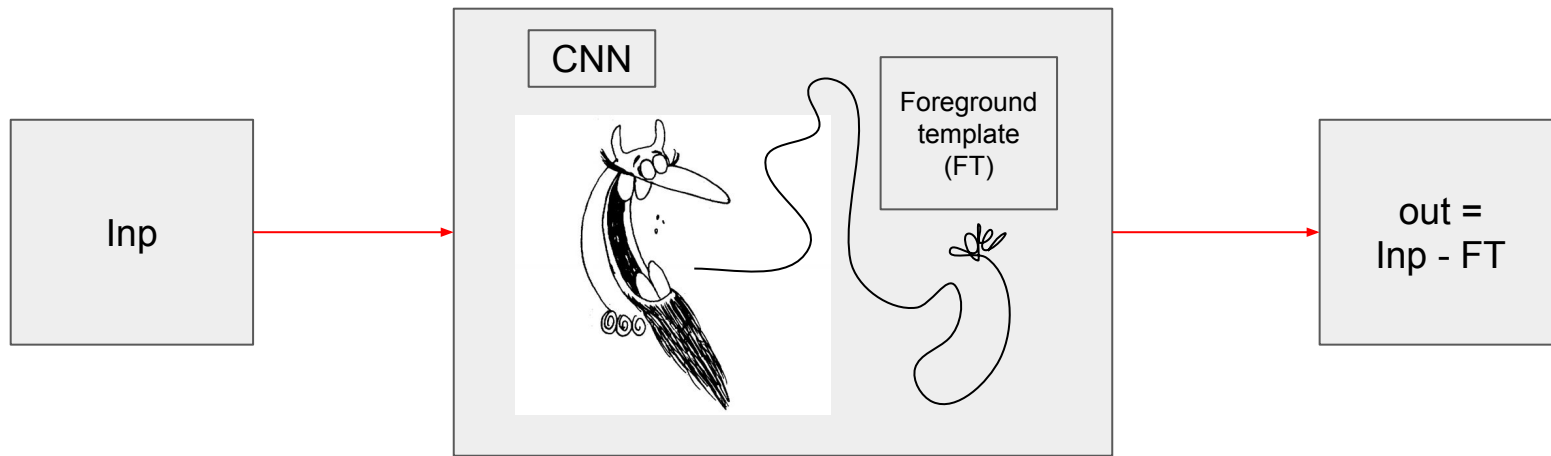
Our current research line



But the principal question of this research line:

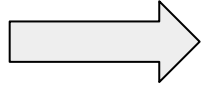
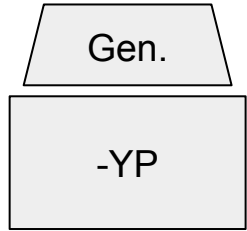
How can we be sure that the CNN is not just subtracting a learned template?

The answer of this question relies in the generalization.

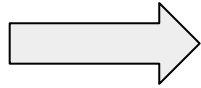
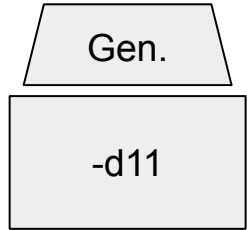


Pictorial representation of Maxwell's demon taken from a lecture given by Esteban Calzetta. I drew the elastic hand.

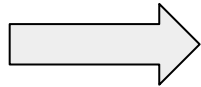
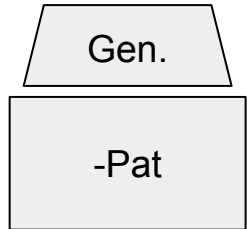
Template based foreground simulation generators



Gaussian fluctuations of the foregrounds parameters for each realization

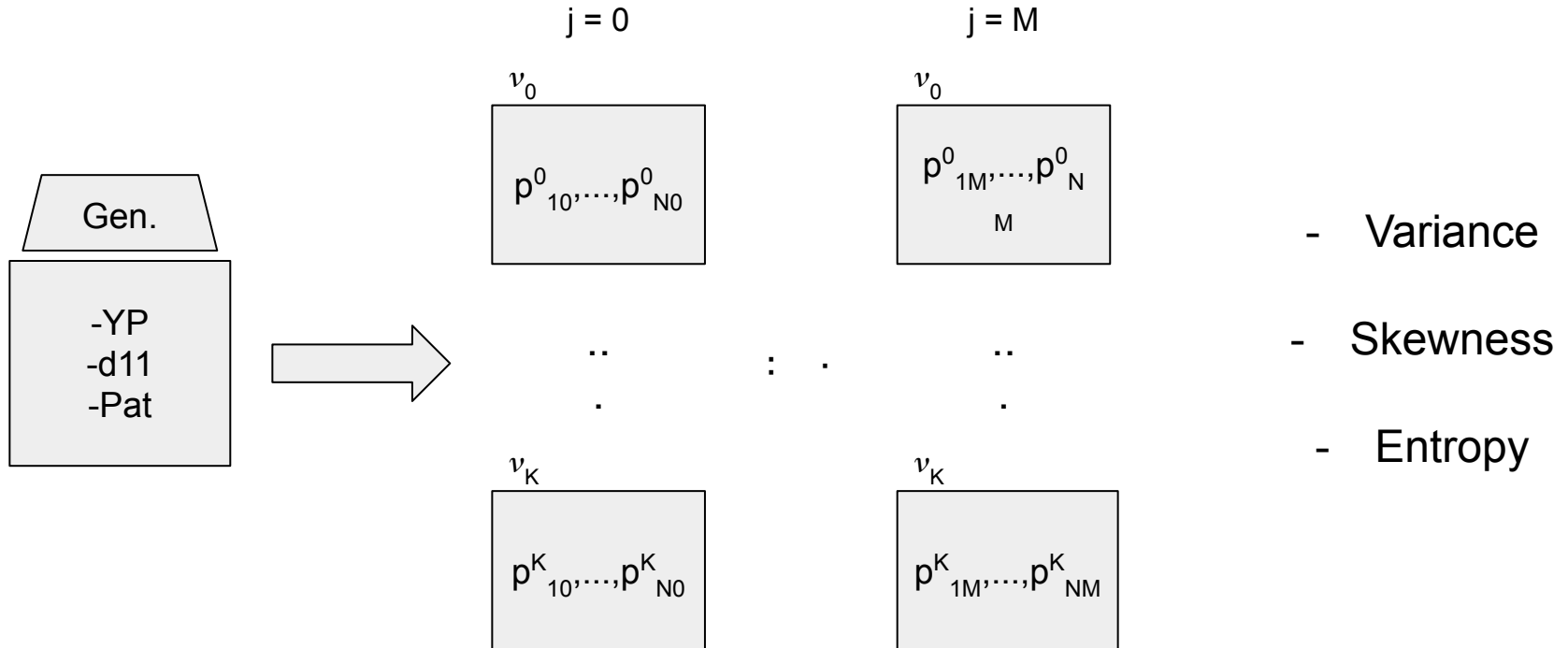


Small scale fluctuations based in a spectra for each realization

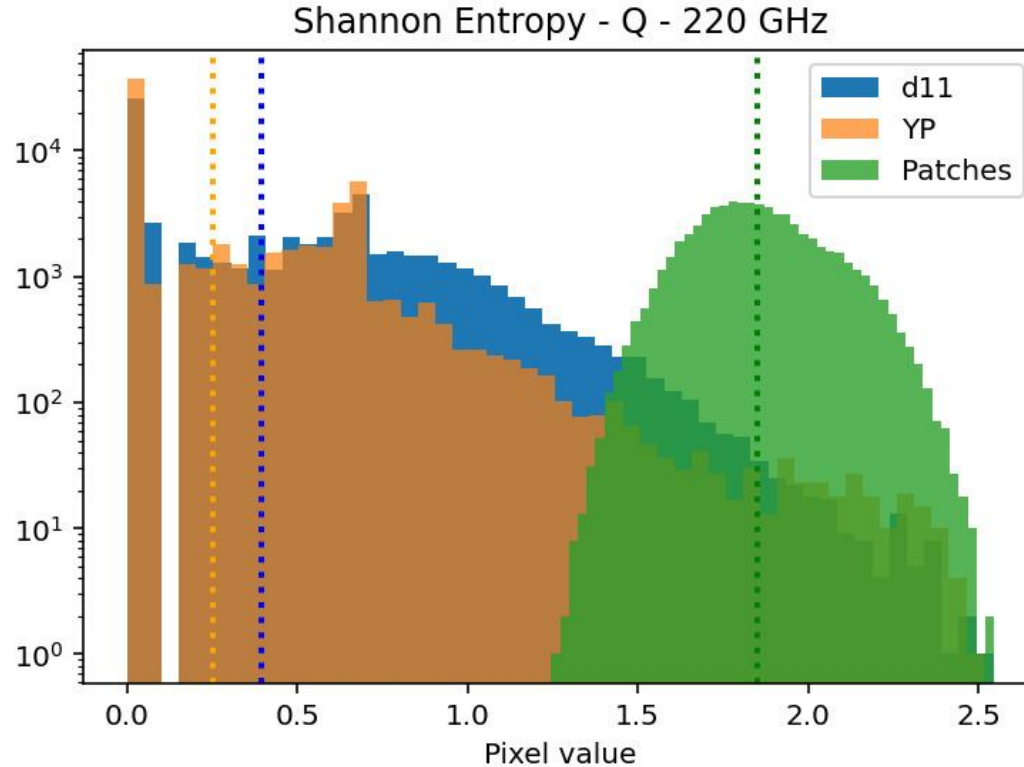


Each realization is a different patch of the sky

Foreground simulation generators

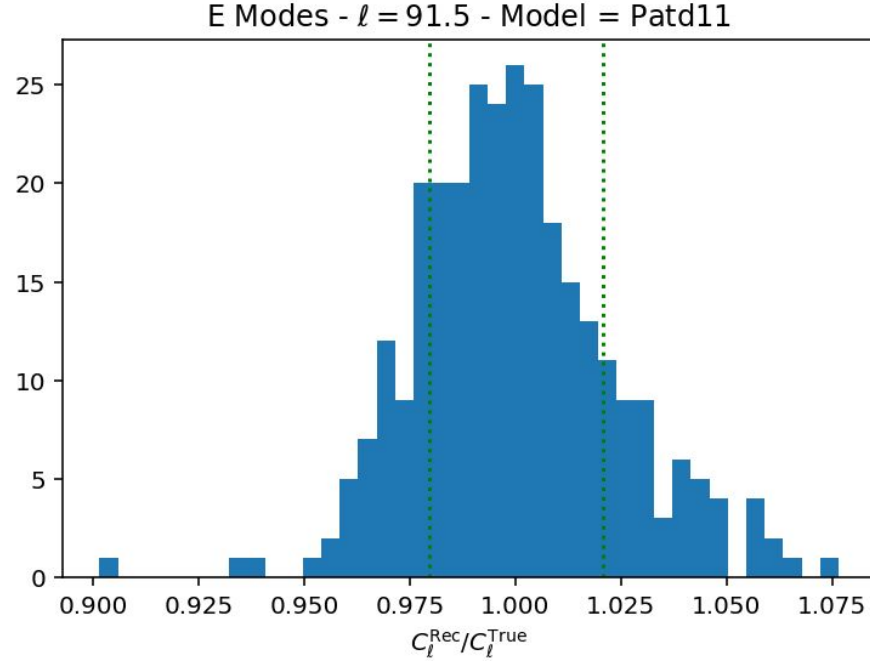


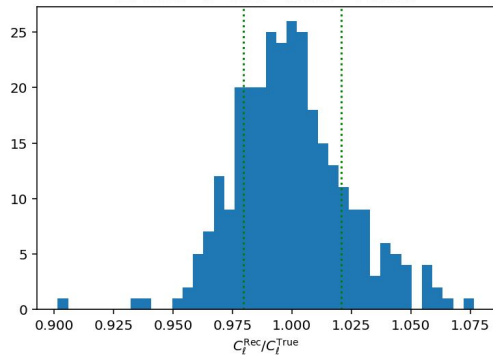
Shannon Entropy for our models



Generalization factor

$$\text{rat}_\ell^{X,i}[\mathbf{M}_1, \mathbf{M}_2] = \frac{C_\ell^{\text{Rec}, X, i}[\mathbf{M}_1, \mathbf{M}_2]}{C_\ell^{\text{True}, X, i}[\mathbf{M}_2]}$$



E Modes - $\ell = 91.5$ - Model = Patd11

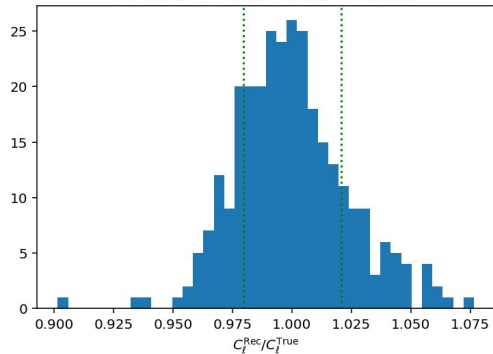
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$$\text{Med}_{\ell}^X[M_1, M_2] = \text{median} \left[\text{rat}_{\ell}^{X,i}[M_1, M_2] \right]_i$$

$$p_{\ell}^{17,X}[M_1, M_2] = \text{percentile} \left[\text{rat}_{\ell}^{X,i}[M_1, M_2], 17\% \right]_i$$

$$p_{\ell}^{83,X}[M_1, M_2] = \text{percentile} \left[\text{rat}_{\ell}^{X,i}[M_1, M_2], 83\% \right]_i$$

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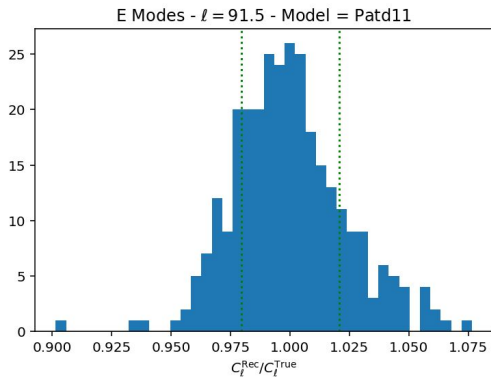
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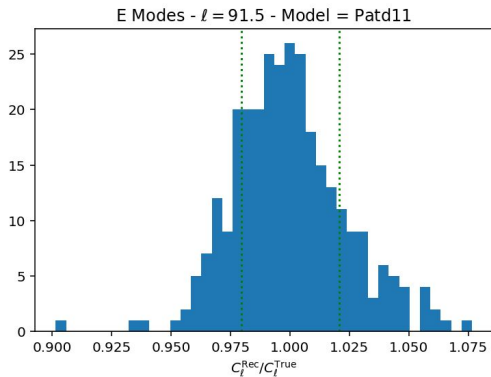
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$$\Delta_\ell^X[\mathbf{M}_1, \mathbf{M}_2] = U_\ell^X[\mathbf{M}_1, \mathbf{M}_2] - L_\ell^X[\mathbf{M}_1, \mathbf{M}_2]$$

$$= p_\ell^{17,X}[\mathbf{M}_1, \mathbf{M}_2] + p_\ell^{83,X}[\mathbf{M}_1, \mathbf{M}_2]$$

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Generalization factor

$$\text{rat}_\ell^{X, i} [M_1, M_2] = \frac{C_\ell^{\text{Rec}, X, i} [M_1, M_2]}{C_\ell^{\text{True}, X, i} [M_2]}$$

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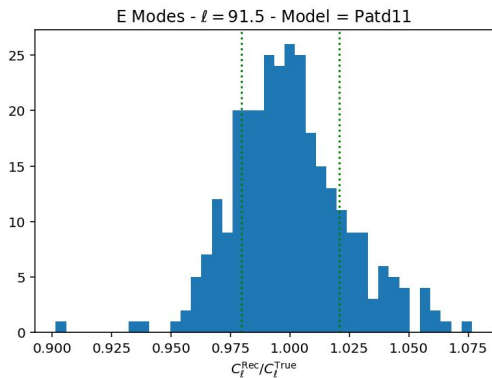
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$$\text{MSE}_\ell^X [M_1, M_2] = \left(\text{Med}_\ell^X [M_1, M_2] \right)^2 + \left(\Delta_\ell^X [M_1, M_2] \right)^2$$



Generalization factor

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$$\mathcal{G}_\ell^X [M_1, M_2; M_3] = \frac{\text{MSE}_\ell^X [M_1, M_3] - \text{MSE}_\ell^X [M_2, M_3]}{\text{MSE}_\ell^X [M_3, M_3]}$$

Generalization factor

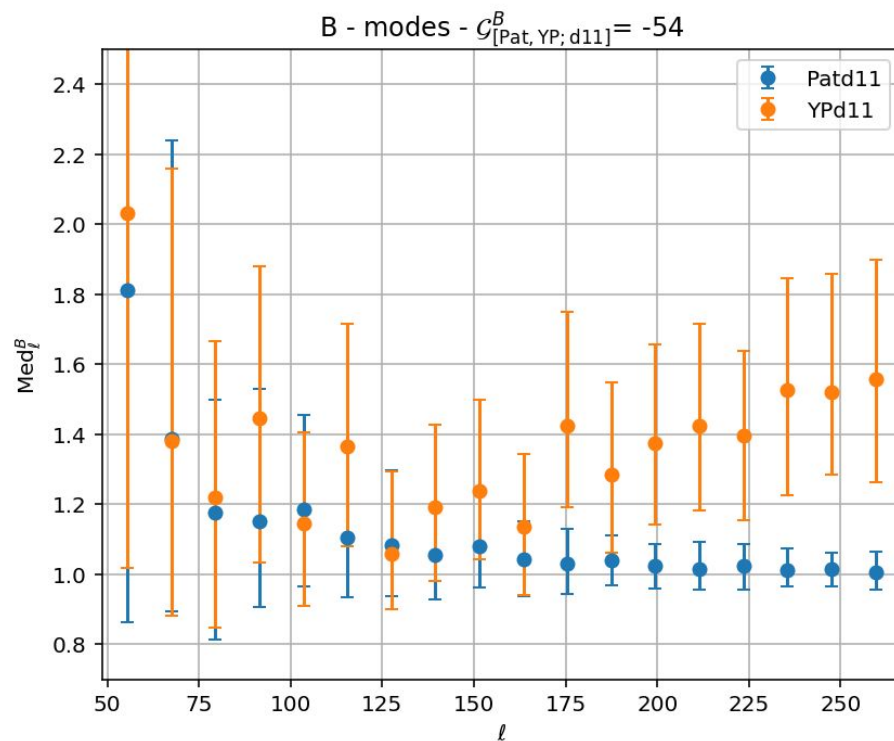
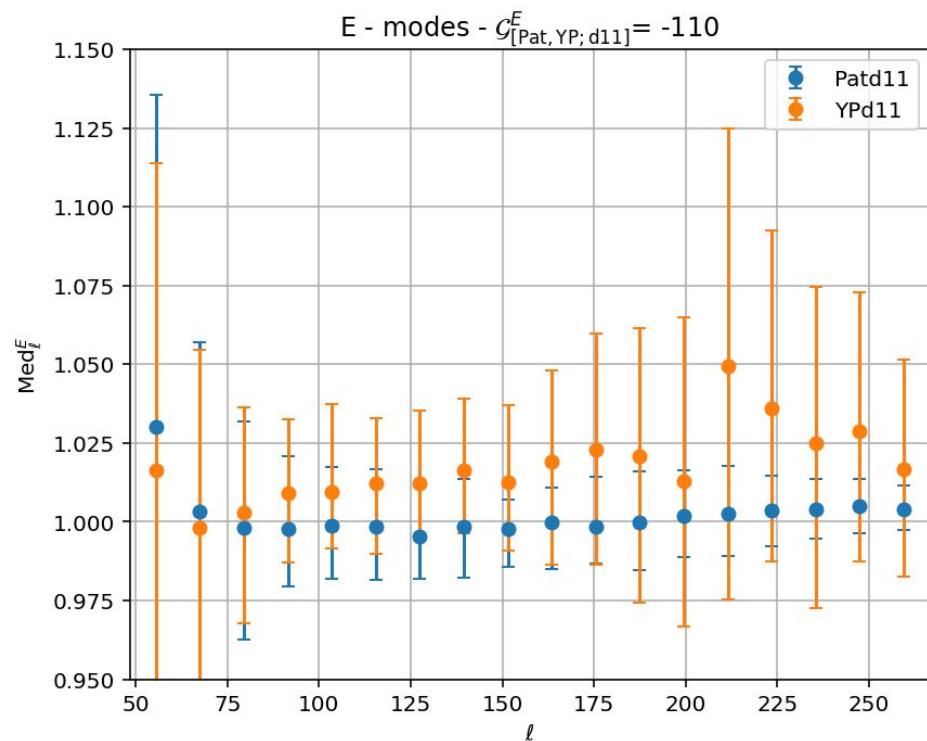
$$\mathcal{G}_\ell^X[M_1, M_2; M_3] = \frac{\text{MSE}_\ell^X[M_1, M_3] - \text{MSE}_\ell^X[M_2, M_3]}{\text{MSE}_\ell^X[M_3, M_3]}$$



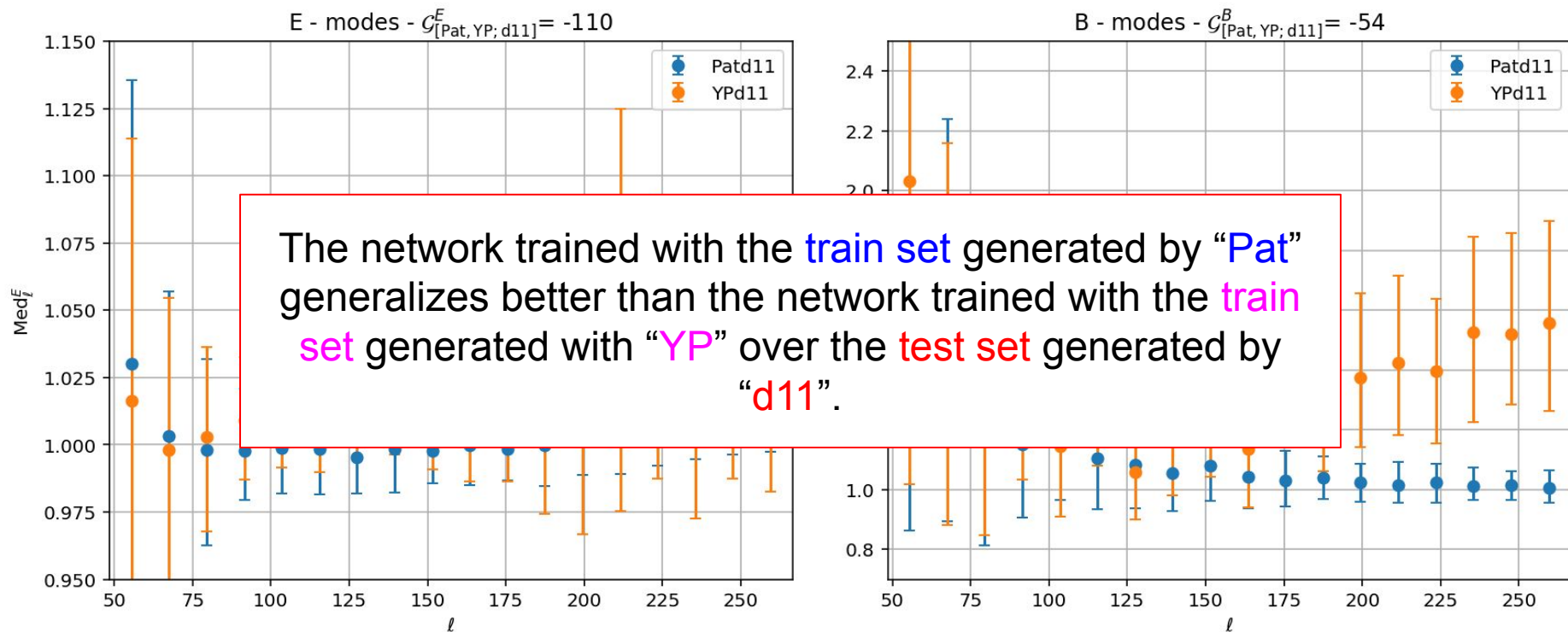
$$\mathcal{G}^{M_1, M_2; M_3} = \langle \mathcal{G}_\ell^X[M_1, M_2; M_3] \rangle_{\ell_{\min} < \ell < \ell_{\max}}$$

- If $\mathcal{G}^{M_1, M_2; M_3} < 0 \rightarrow$ M1 generalizes better than M2 on M3
- If $\mathcal{G}^{M_1, M_2; M_3} > 0 \rightarrow$ M2 generalizes better than M1 on M3

Preliminary result: [Pat, YP; d11]



Preliminary result: [Pat, YP; d11]



Thank you for the attention!