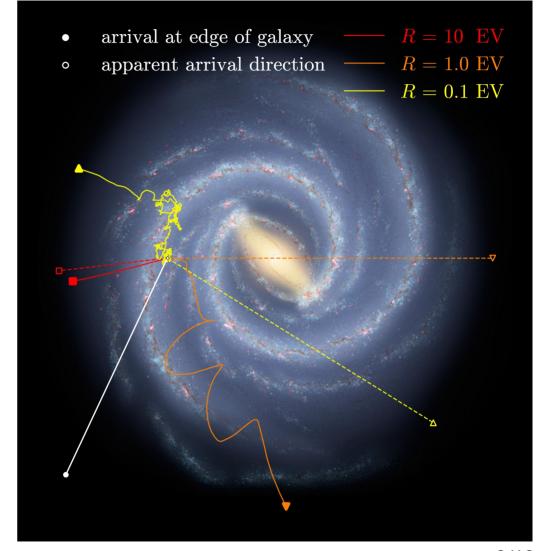
HE Photon search at the Pierre Auger Observatory

Paul Filip

(Part III)

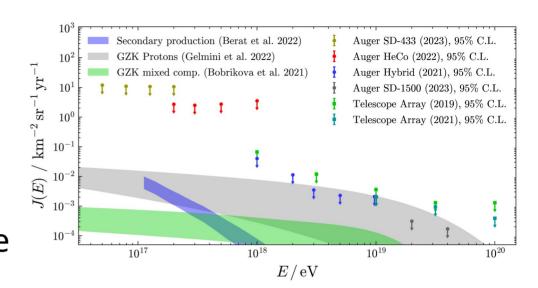
Why care about photons?

- Cosmic rays are charged
 - Deflected by Lorentz force
 - Diffusive + ballistic trajectory
 - Direction essentially random
- Photons are uncharged
 - Not affected by magn. fields
 - Travel in straight lines
 - Point directly to their source



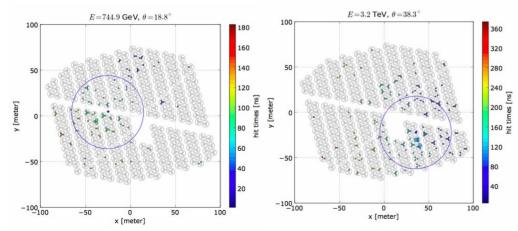
What do we know so far?

- Potential for the direct observation of the universes most energetic processes
- Photons above $10^{17}\,\mathrm{eV}$ very rare

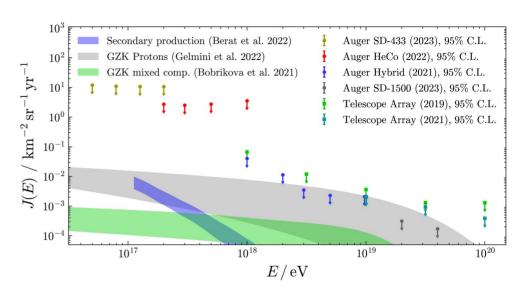


What do we know so far?

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https://hawc-observatory.org/observatory/ghsep.php



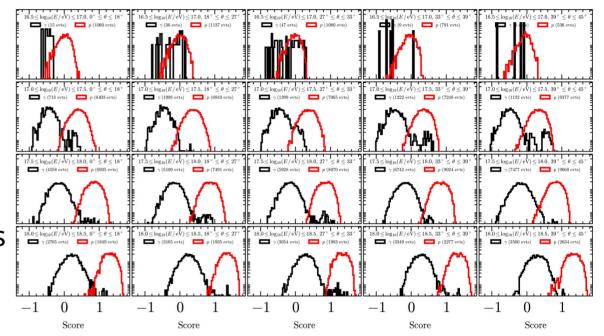
- Need to discriminate between photon/hadron showers
- Difficult on an event-by-event basis → probabilistic analysis over many showers

Discrimination between photons and hadrons

 Use a score to judge the "photonness" of a shower

-
$$M_b = \log \left(\sum_i \frac{\rho_i}{\rho_{\text{Pr}}(E, \theta)} \times \left(\frac{r_i}{r_{\text{Pr}}} \right)^b \right)$$

- M_b not only existing score
- Calculate from simulations

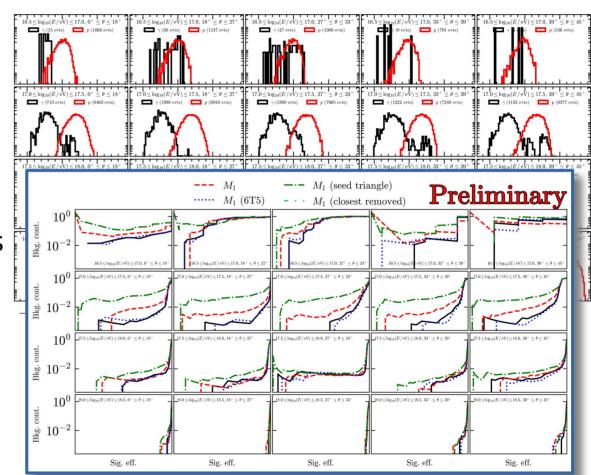


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- Calculate from simulations
- Decide if shower is p/γ based on cut in score
 - Calculate signal efficiency and bkg. contamination



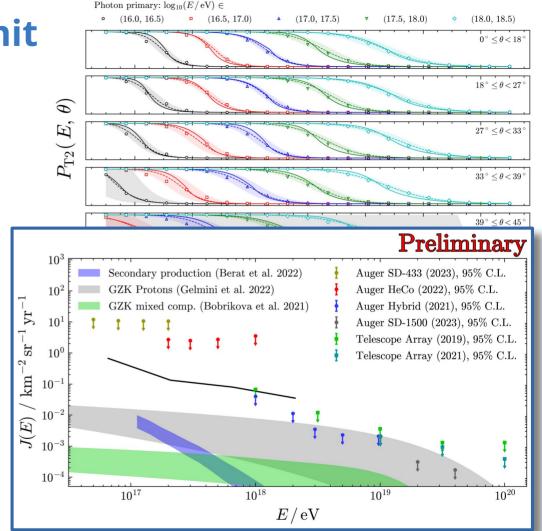
11/11/2025 paul.filip@kit.edu 6/12

Calculation of upper flux limit

Calculate upper flux limit as:

$$J_{\text{max}} = \frac{n(\rho_{\gamma})}{e \,\epsilon_{\text{T3}}(E, \,\theta) \,\rho_{\gamma}(E, \,\theta)}$$

- Where $\epsilon_{\rm T3}$ = detection chance
- ρ_{γ} = detected photon fraction
- $n(\rho_{\gamma})$ = confidence limits
- Decrease current world best limits by a factor of 10-100

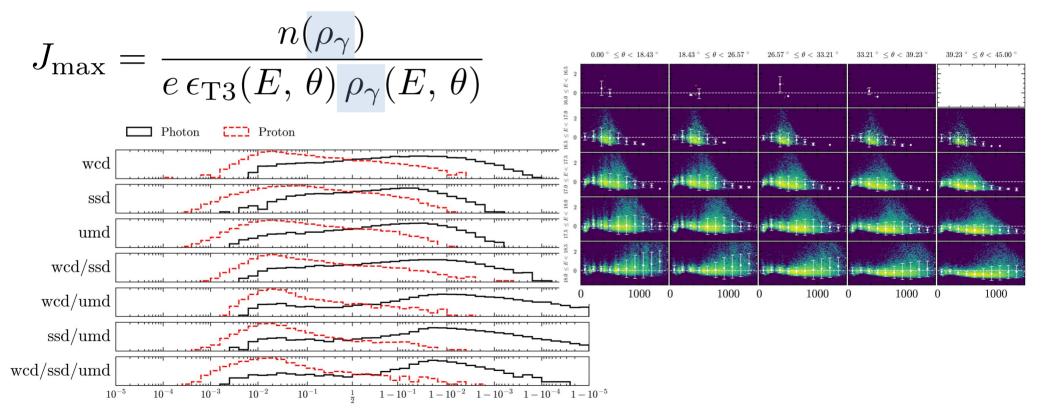


Calculate upper flux limit as:

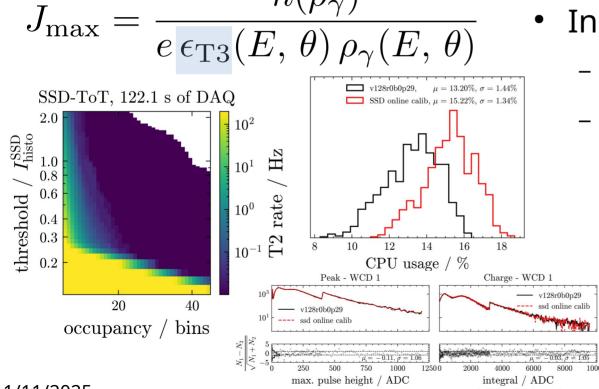
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Calculate upper flux limit as:

Increase discrimination power

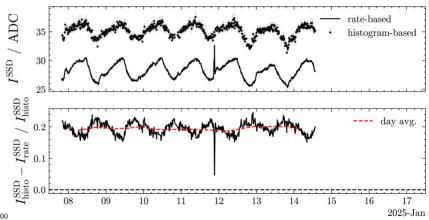


Calculate upper flux limit as:



Increase discrimination power

- Increase trigger efficiency
 - new detectors → new triggers?
 - new SSD calibration algorithm



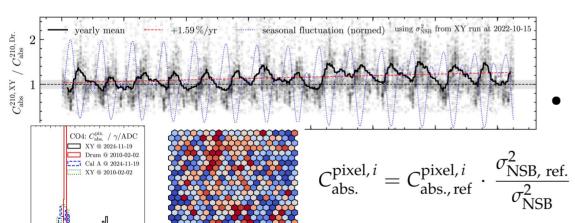
11/11/2025

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10/12

Calculate upper flux limit as:

$$J_{\max} = rac{n(
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(XY / Drum) @ 2010-02-02

- Increase discrimination power
- Increase trigger efficiency
 - new detectors → new triggers?
 - new SSD calibration algorithm
 - Revisit Augers energy scale
 - Evidences of telescope aging
 - FD energy changes w/ time
 - Attempt to correct this

Backup