



PIERRE  
AUGER  
OBSERVATORY



# Depth of maximum of air-shower profiles: testing the compatibility of the measurements at the Pierre Auger Observatory and the Telescope Array

**Auger:** Jose Bellido, Ralph Engel, Vitor de Souza, Eric Mayotte, Olena Tkachenko, Michael Unger, [Alexey Yushkov](#)



**Telescope Array:** John Belz, Douglas Bergman, Toshihiro Fujii, Zane Gerber, Daisuke Ikeda, Jihyun Kim, Yoshiki Tsunesada

for the Pierre Auger and Telescope Array Collaborations



20/11/2024

# Data sets from the Fluorescence Detectors

## Pierre Auger

Surface Detector: 3000 km<sup>2</sup>

13 years of data (12/2004 – 12/2017)

12773 events<sup>†</sup>

[Auger, ICRC (2019), PoS 482]

## Telescope Array

Surface Detector: 700 km<sup>2</sup>

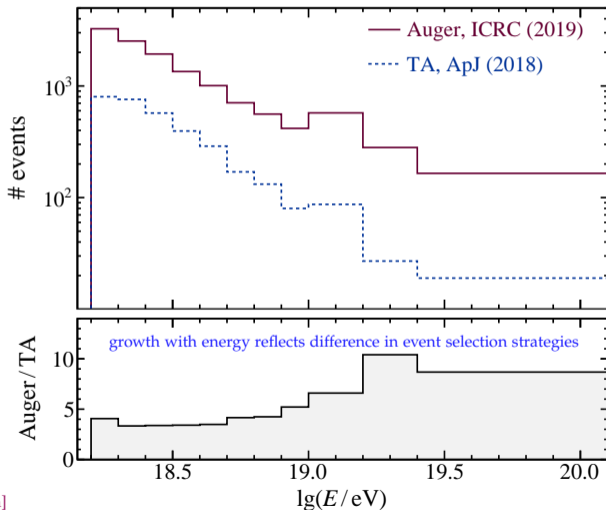
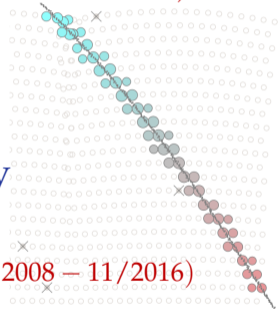
8.5 years of data (05/2008 – 11/2016)

3330 events

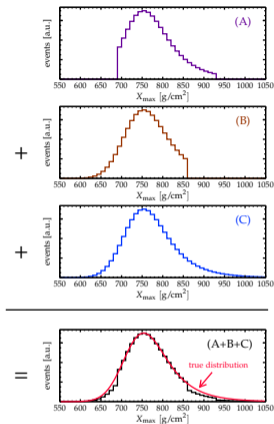
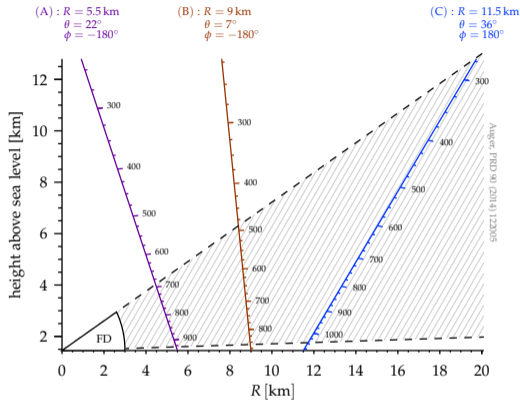
[Ap] 858 (2018) 76]

<sup>†</sup> Auger (12/2004 – 12/2021): 17605 events (+38%) [T. Fitoussi, this session]

common energy range:  $\lg(E/\text{eV}) > 18.2$ , TA energy binning



# Event selection strategies: detector acceptance



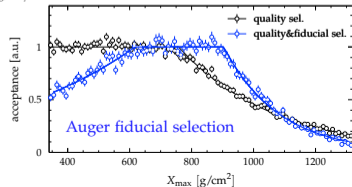
Taking all geometries:  
smaller relative acceptance  
to shallow and deep events

Pierre Auger

select geometries similar to (C), efficiency  $\approx$  40%  $\lg(E/\text{eV}) = 18.2$   
60%  $\lg(E/\text{eV}) = 19.2$

Telescope Array

keep all geometries to maximize statistics



# $X_{\max}$ moments measured at Auger

## Event selection

◇ fiducial selection: minimal acceptance biases on tails of  $X_{\max}$  distributions

$$\langle X_{\max} \rangle$$

◇ corrected for residual acceptance and reconstruction biases

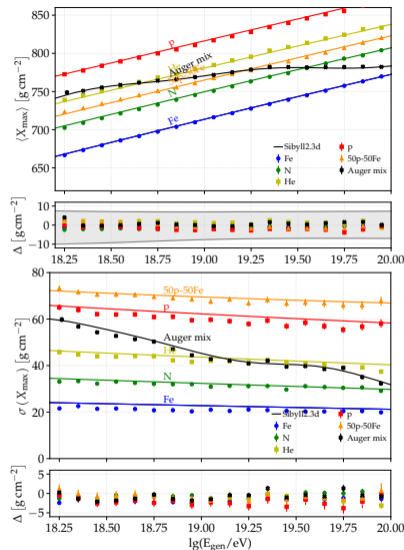
$$\sigma(X_{\max})$$

◇ detector resolution (including atmosphere) is subtracted

$\langle X_{\max} \rangle, \sigma(X_{\max})$  are unbiased, directly comparable to predictions of air-shower simulation codes

## $X_{\max}$ distributions

◇ simulations should be folded with the detector effects  $\otimes$  Auger



lines — air-shower simulations

points — simulations after full analysis chain

Auger (2024), T. Froust, this session

## Event selection

◇ fiducial selection: minimal acceptance biases on tails of  $X_{\max}$  distributions

$\langle X_{\max} \rangle$

◇ corrected for residual acceptance and reconstruction biases

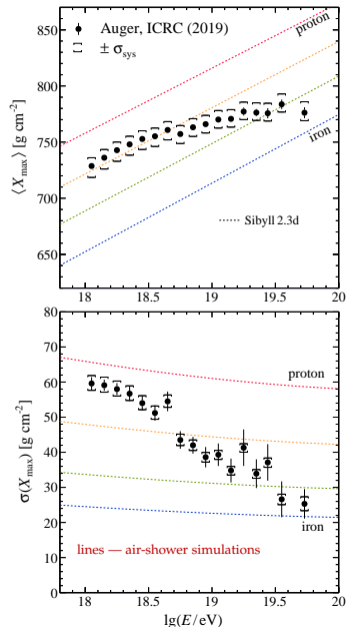
$\sigma(X_{\max})$

◇ detector resolution (including atmosphere) is subtracted

$\langle X_{\max} \rangle, \sigma(X_{\max})$  are unbiased, directly comparable to predictions of air-shower simulation codes

$X_{\max}$  distributions

◇ simulations should be folded with the detector effects ⊗ Auger



## Event selection

◇ no fiducial selection, acceptance biases on tails of  $X_{\max}$  distributions

$\langle X_{\max} \rangle$

◇ acceptance biases are mostly seen for protons

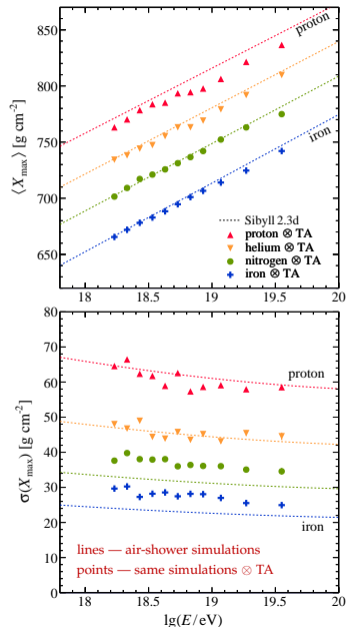
$\sigma(X_{\max})$

◇ detector resolution effect is visible for nitrogen and iron

$\langle X_{\max} \rangle, \sigma(X_{\max})$  are not corrected for experimental biases and are not directly comparable to predictions of air-shower simulation codes

## $X_{\max}$ distributions

◇ simulations should be folded with the detector effects  $\otimes$  TA



## Event selection

◇ no fiducial selection, acceptance biases on tails of  $X_{\max}$  distributions

$$\langle X_{\max} \rangle$$

◇ acceptance biases are mostly seen for protons

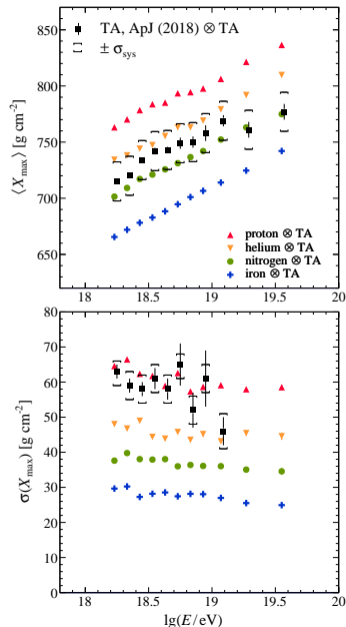
$$\sigma(X_{\max})$$

◇ detector resolution effect is visible for nitrogen and iron

$\langle X_{\max} \rangle, \sigma(X_{\max})$  are not corrected for experimental biases and are not directly comparable to predictions of air-shower simulation codes

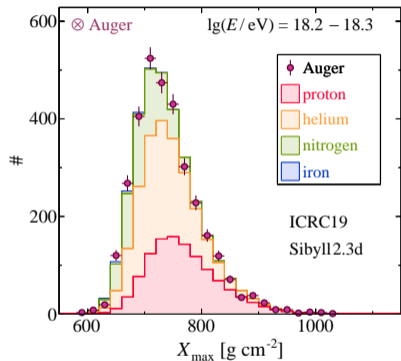
## $X_{\max}$ distributions

◇ simulations should be folded with the detector effects  $\otimes$  TA

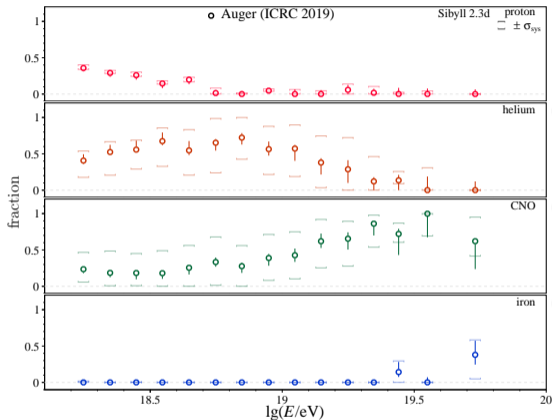


# Method to transfer Auger data into TA detector

Combine (p, He, N, Fe) of Sibyll2.3d to fit Auger  $X_{\max}$  distributions



Energy evolution of nuclear fractions in fits of the Auger  $X_{\max}$  distributions



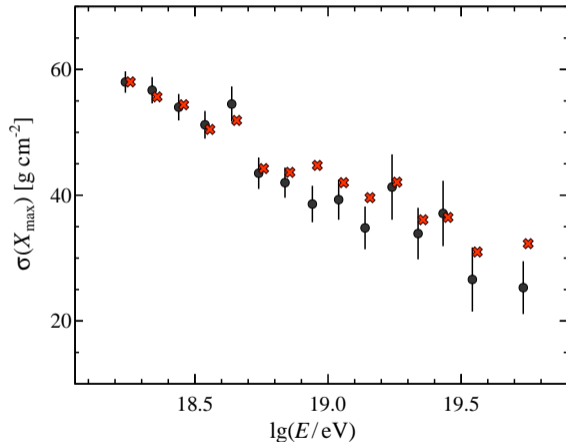
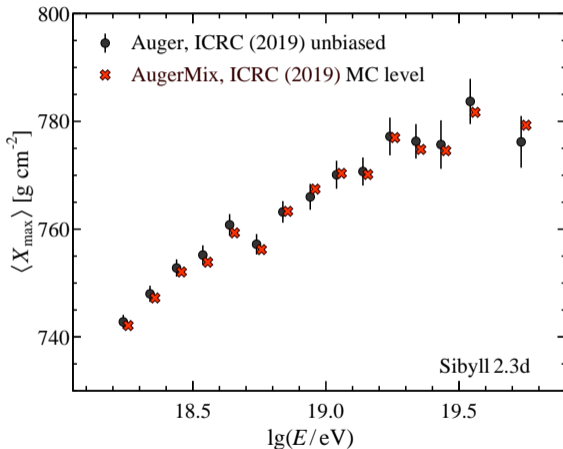
Use as a proxy simulated  $X_{\max}$  mixes (AugerMix) and process them with the TA machinery

Result: AugerMix  $\otimes$  TA — representation of Auger  $X_{\max}$  distributions folded with the TA detector and analysis effects



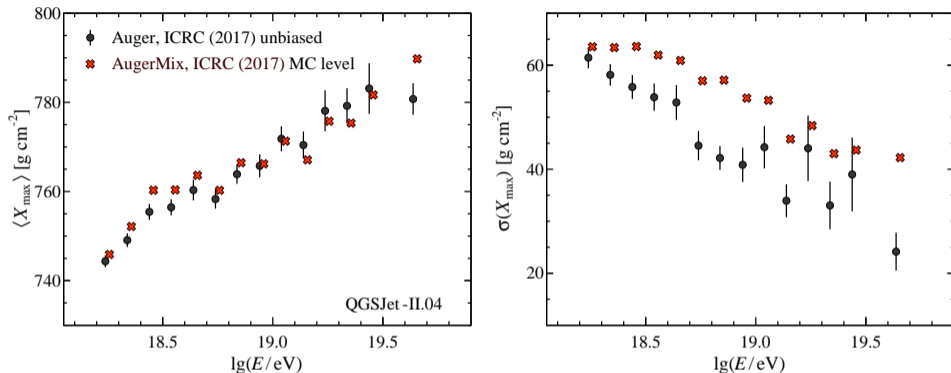
# $X_{\max}$ moments in AugerMix and Auger data

Good description with AugerMix of  $\langle X_{\max} \rangle$  and  $\sigma(X_{\max})$  measured at Auger



MC level —  $X_{\max}$  moments directly from Monte-Carlo air-shower simulation codes (no detector effects)

# $X_{\max}$ moments in AugerMix and Auger data: why QGSJet-II.04 can not be used



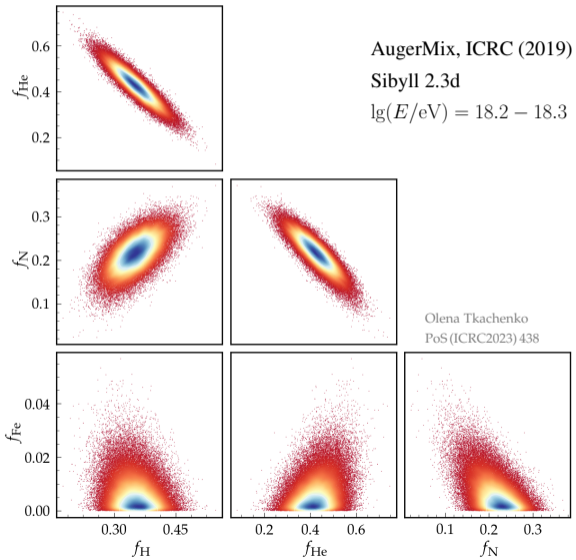
- ◇ width of the observed  $X_{\max}$  distributions is not reproduced well
- ◇ p-values in the fits Auger data are  $\approx 0.01$  for  $\lg(E/\text{eV}) = 17.8 - 19.2$

**QGSJet-II.04  $\langle X_{\max} \rangle$  predictions are strongly disfavored by the Auger data**

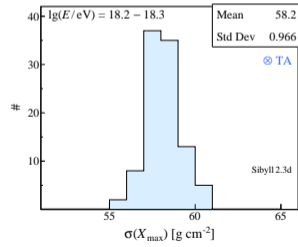
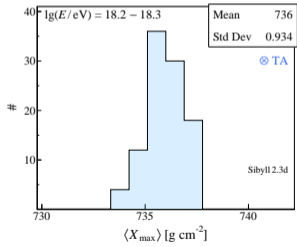
[most recent Auger papers: [PRD 109 \(2024\) 102001](#); DNN SD  $X_{\max}$  with  $> 10 \times$  FD statistics, accepted by [PRD](#) and [PRL](#)]

# Estimation of statistical uncertainties on $X_{\max}$ moments $\otimes$ TA

Markov Chain Monte Carlo fits: posterior distributions of fractions



- ◇ pick randomly 100 mixes from posterior distributions
- ◇ prepare 100 AugerMix samples  $\otimes$  TA
- ◇ use  $1\sigma$  of their  $\langle X_{\max} \rangle$  and  $\sigma(X_{\max})$  distributions as an estimate of statistical uncertainties

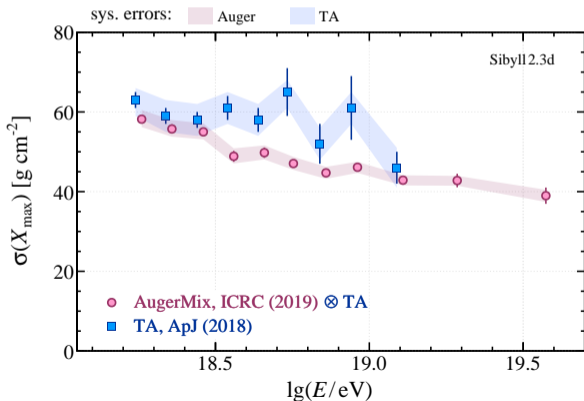
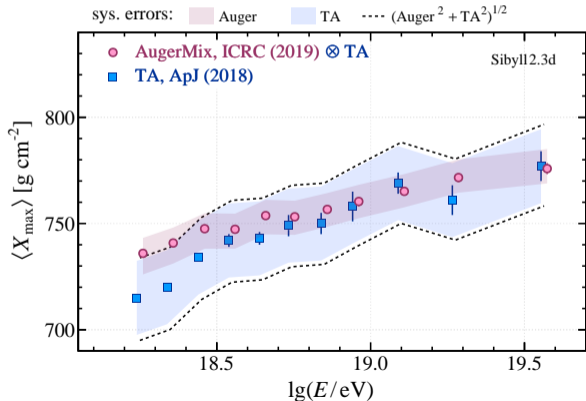


# Comparison of the $X_{\max}$ moments measured at TA and Auger $\otimes$ TA

$\langle X_{\max} \rangle$  — agreement withing statistical and systematic uncertainties, in particular for  $\lg(E/\text{eV}) > 18.5$

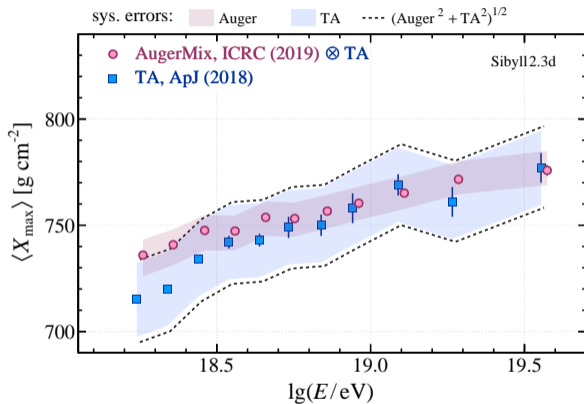
$\sigma(X_{\max})$  — larger values in TA for  $\lg(E/\text{eV}) = 18.5 - 19.0$ , possible reasons:

- ◇ constant aerosol profiles used in TA increase  $\sigma(X_{\max})$  by  $18.9 \text{ g cm}^{-2}$  (in quadrature) [ApJ 858 (2018) 76]
- ◇ a few deep events in data can increase  $\sigma(X_{\max})$  significantly (see  $X_{\max}$  distributions in next slides)



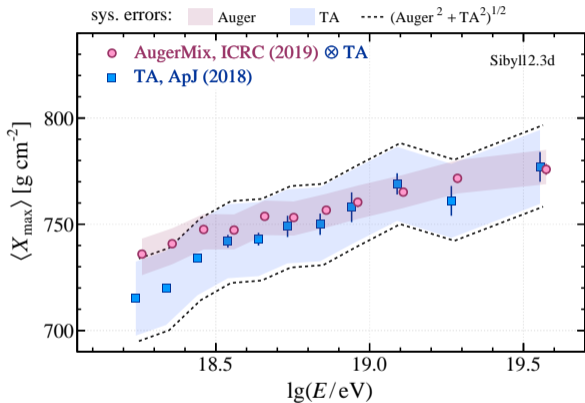
# Discrepancy in $\langle X_{\max} \rangle$ at lower energies

TA  $\langle X_{\max} \rangle$  is shallower at  $\lg(E/\text{eV}) < 18.7$ ; difference is  $\sim 20 \text{ g cm}^{-2}$  for  $\lg(E/\text{eV}) = 18.2 - 18.4$



# Discrepancy in $\langle X_{\max} \rangle$ at lower energies

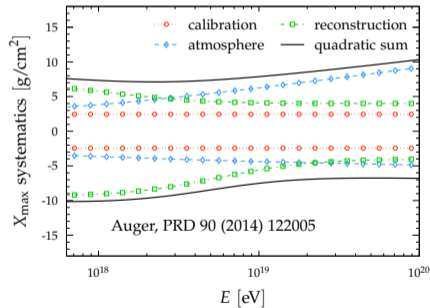
TA  $\langle X_{\max} \rangle$  is shallower at  $\lg(E/\text{eV}) < 18.7$ ; difference is  $\sim 20 \text{ g cm}^{-2}$  for  $\lg(E/\text{eV}) = 18.2 - 18.4$



Far beyond  $\sigma_{\text{sys}}(X_{\max})$  energy dependence

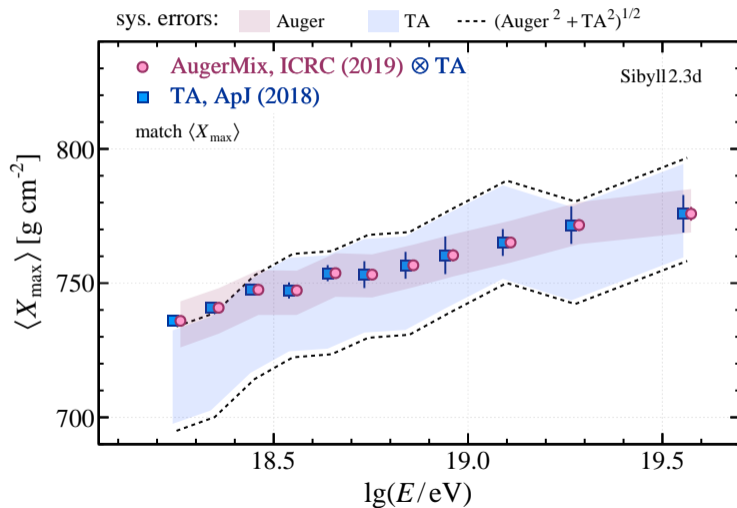
TA: constant  $\sigma_{\text{sys}}(X_{\max}) = 17.4 \text{ g cm}^{-2}$

Auger:  $\lesssim 5 \text{ g cm}^{-2}$  change in this energy range



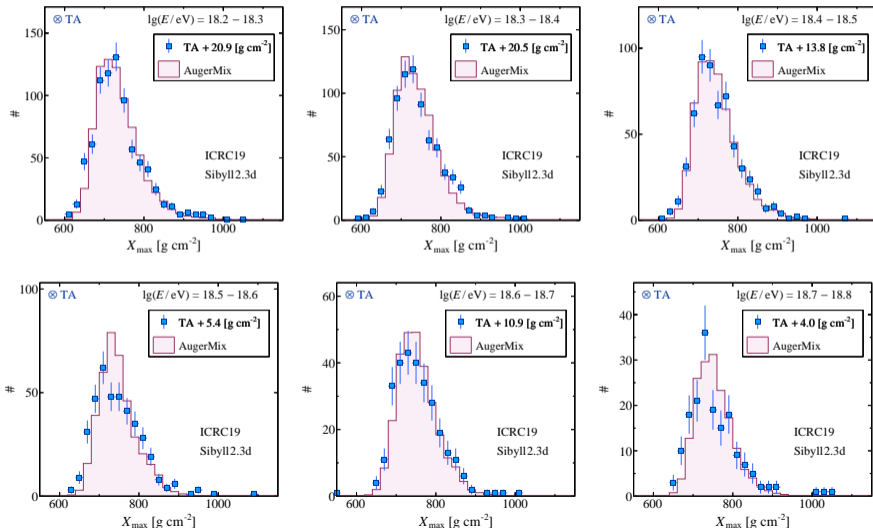
# Compatibility of $X_{\max}$ distributions

To compare shapes, we align  $\langle X_{\max} \rangle$  of AugerMix and TA  $X_{\max}$  distributions



# Comparison of TA and AugerMix $X_{\max}$ distributions

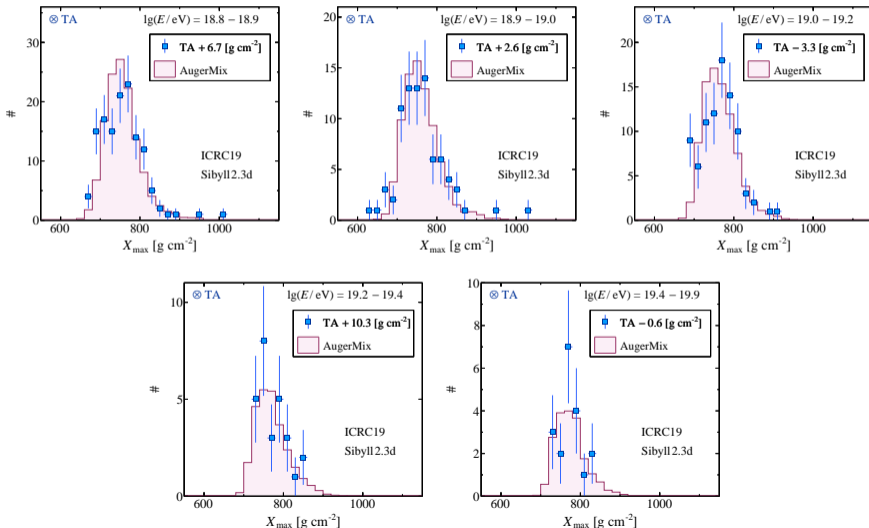
Visually, good agreement in most of the energy bins





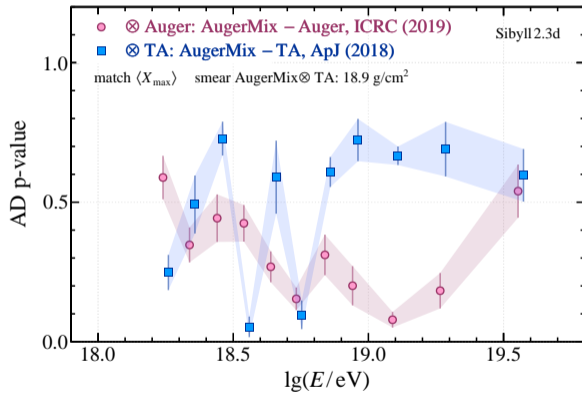
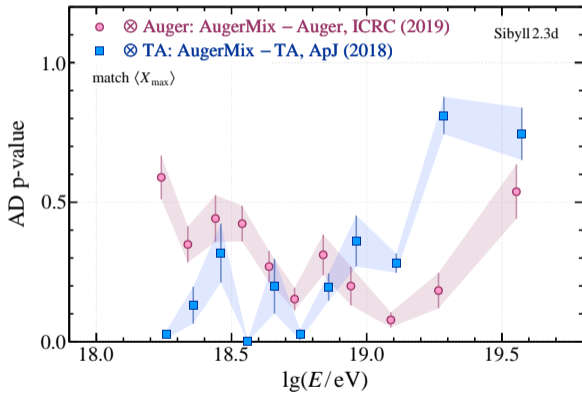
# Comparison of TA and AugerMix $X_{\max}$ distributions

Visually, good agreement in most of the energy bins



# Compatibility of distributions: Anderson-Darling test

Good agreement in most energy bins



Constant aerosols: AugerMix ⊗ TA smeared by 18.9 g cm<sup>-2</sup>

$\langle X_{\max} \rangle$  and  $\sigma(X_{\max})$  generally agree within statistical and systematic uncertainties

◇ origin of the energy trend in  $\langle X_{\max}^{\text{Auger}} \rangle - \langle X_{\max}^{\text{TA}} \rangle$  at lower energies is not clear

◇  $\sigma(X_{\max}^{\text{TA}}) > \sigma(X_{\max}^{\text{Auger}})$  — partly can be attributed to the use of a static atmosphere model at TA

Shapes of  $X_{\max}$  distributions agree well in most energy bins

Auger and TA measurements are compatible at the current level  
of statistics and understanding of systematics

Draft of the joint paper is under collaborations review