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## Individual particles at the highest energies:

- Events at the highest energies have the potential to be the most constraining, higher energies suggest closer horizons and trajectories that are less impacted by magnetic fields.
- Amaterasu is the second highest energy particle ever detected, E<sub>det</sub> = (244 ± 29 (stat.)<sup>+51</sup><sub>-76</sub>(syst.)) EeV [1], its arrival direction aligns with the Local Void.





### Methods:

- We use CRPropa3 [2] to simulate the propagation of UHECRs, including energy losses and magnetic fields.
- We make two analyses, one with the nominal energy,  $E_{nom}$  and one where we include the lower bound of the systematic uncertainty,  $E_{low}$ .
- We use Approximate Bayesian Computation, a simulation based inference method, to estimate the posterior distribution of our six free parameters: B<sub>rms</sub>, L<sub>c</sub> of the extra-Galactic magnetic field and D<sub>src</sub>, E<sub>src</sub> and (l, b)<sub>src</sub>.
- ▶ We define a deviation  $d = |X_{sim} X_{exp}|$  and a tolerance  $e = 3\sigma$ , events that satisfy  $|E_{sim} E_{exp}| \le 3\sigma$  and  $|dir_{sim} dir_{exp}| \le 3\sigma$  are accepted.

#### Distance (Mpc)

#### Results:

- We use four galaxy catalogs [3,4,5,8] including star-burst galaxies, active galactic nuclei and regular galaxies to check for overlaps.
- We find that the source should be at  $D_{src} \le 12$  Mpc for  $E_{nom}$  and  $D_{src} \le 15$  Mpc for  $E_{low}$ .
- Our results are dependent on the Galactic magnetic field model we use, JF12<sub>[7]</sub> and UF23<sub>[10]</sub>.

### Conclusion:

- Our results are consistent with region of space beyond the Local Void, especially when considering the systematic uncertainty and the arrival compositon.
- We find multiple possible sources, including several starburst galaxies, like M82.

#### References:

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