







Calibration of the Upgraded Surface Detector Stations of the Pierre Auger Observatory

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Alexander Streich, David Schmidt, Darko Veberič, Markus Roth, Ralph Engel

AugerPrime Surface Detector

What is **new** for the Surface Detector stations?

- Electronics boards HIRSAP 2019
- Small PMT inside water-Cherenkov detectors (WCD)
- Scintillator detectors (SSD):
 - Validation tests and comparisons with simulations HIRSAP 2018
 - Calibration and event data analysis HIRSAP 2020
- Radio antennas





SD Data Processing

Why is calibration important?



\rightarrow Optimization and adaptation of calibration procedure

Current Calibration of WCD and SSD

- Muon hump search by determination of left and right "shoulders" around first maximum
- Check if online value within the shoulder limits
- If fit algorithm fails \rightarrow online charge value used
- Analytic quadratic fit of bins in-between shoulders

SD event: 59432748, Old station: 1113, PMT: WCD3









Current Calibration of WCD and SSD

- Limited flexibility due to several fixed variables (e.g., start values, thresholds)
- WCD: Number of successful fits decreases (change of charge histogram shapes)
- SSD: Filler module for online estimates → incorrect charge calibration







SD event: 59437611, SSD PPA station: 701, PMT: SSD

New Calibration of WCD and SSD

- Search for hump and dip positions (both extremes necessary for successful calibration)
- Calculate 3-bin-density at each bin position: $d_i = \frac{1}{3} \left(\frac{1}{\Delta_{i-1}} N_{i-1} + \frac{1}{\Delta_i} N_i + \frac{1}{\Delta_{i+1}} N_{i+1} \right)$
- Left and right limits around extremes with relative thresholds (10 σ , 8 σ , 5 σ)
- Analytic quadratic fit of bins in-between limits (for dip optional)



New Calibration of WCD

• New algorithm only successful when hump and dip determined

 \rightarrow Higher quality, but lower number of passing histograms



New Calibration of SSD

• No check for online estimate in-between hump limits

 \rightarrow Incorrect filler module estimate not used



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Long-term Performance

- Calibration depends on shape of histograms which changes with time
- Prediction of future performance
 - \rightarrow reaching limits of new calibration





ddir



Long-term Performance

• Ratio of hump position (charge) and hump width

 \rightarrow Light yield estimate: $N_{\rm eff} \approx \left(\frac{Q_{\rm hump}}{W_{\rm hump}}\right)^2$





Air Shower Reconstruction

- Comparison: current to new calibration algorithm
- Rejection of PMTs possible \rightarrow rejection of stations possible \rightarrow stations in events can differ



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SSD Pre-production Array

- Data acquisition since March 2019
- SSD PMT connected to non-upgraded electronics (one WCD PMT disconnected)



SSD Pre-production Array

- Data up to November 2020
- Events with zenith angle $\theta < 60^{\circ}$



Look into the Future

- Production of **1518** SSDs completed
- Ongoing deployment of AugerPrime components
- Very near future: UUB pre-production array (start with 31 new electronics boards)



Look into the Future

- Challenge: Calibration algorithm that suits different types of electronics and PMTs
- Modified algorithm more flexible with higher number of successful fits
- Very near future: Adaptations to UUB PPA data

The glorious multi-hybrid detection and data analysis starts now! 613

Thank You!



Rejected PMTs

- PMTs rejected by CDAS or during event merging (quality flags from monitoring)
- Additional rejection by certain online values (only available for UB WCD PMTs)



Non-upgraded stations: PMT rejected

Calibration of WCD and SSD

- Pre-selection: Rejection of anomalous histograms before fit procedure
- Check of total number of entries and histogram entropy







Calibration of WCD and SSD

- Post-selection: Quality check of fit results
- Requirement of minimum hump/dip charge ratio → rejection of fits of fluctuations



Air Shower Reconstruction

• Current calibration algorithm:



Air Shower Reconstruction

• New calibration algorithm:

